## Dissertation

on

# STUDIES IN COSMIC RAYS

"The Study of the Daily Variation of Locally Produced Neutrons"

and

"An all Copper Metal Systèm for Purification of Noxious and Inflammable Gases"

presented

by

SATYA PRAKASH

for the degree of

DOCTOR OF PHILOSOPHY

of the

Gujarat University

043

B1927

October 1958

Physical Research Laboratory
Ahmedabad.

#### STATEMENT.

- 1. The daily variation and the daily mean intensity of local neutron measured with a neutron monitor ( $N_k$ ) at the equatorial mountain station of Kodaikanal ( $\lambda_m = 1^\circ$ , altitude 7688 ft.) during the period November 1956 to September 1957, has been studied by the author. The daily variation and the daily mean intensity were also studied in relation to the geomagnetic disturbance, as indicated by  $C_p$ . The variance has been taken as an index of the variability of intensity during successive bihourly intervals of 12 bihours, and has been used for defining days with high and low anisotropy. The amplitude and phase of the diurnal and the semidiurnal components of the daily variation for local neutrons are also compared with those of the daily variation of the meson component measured with telescopes ( $M_k$ ) having a semiangle of opening 22° in the E-W plane and 37° in N-S plane.
- 2. The study indicates the following:
  - a. From the study of the histogram of days grouped according to the magnitude and sign of the percent deviations, ΔI of the daily mean intensity from the mean intensity of the whole period, it can be concluded that decreases of daily mean intensity are much larger in amplitude than increases.
  - b. A study of the histogram of days grouped according to the magnitude of variance  $\Sigma^2$  of each individual day reveals that there are a large number of days which show a variance much larger than can be expected from Poisson's distribution.

- c. From the study of occurrence of significant bihourly deviations at different bihours and the study of the mean daily variation for days on which a significant bihourly deviation at a particular bihour occurs, it is found that there is a region of broad maximum at 10, 12, 14 hours and a region of sharp maximum at 0 hour. There are two regions of minima, one at 6 hour (a sharp minimum) and the other at 18, 20, 22 hours (a broad minimum). The maximum at 12 and 14 hours and the minimum at 18 and 20 hours are the most prominent features of the daily variation.
- d. The study of the mean daily variation I and the histogram of days grouped according to the times of maxima of the diurnal and the semidiurnal components reveal that with decrease of intensity the time of maximum of the diurnal component shifts to earlier hours whereas the second harmonic is not significantly affected.

Such a study for groups according to the value of  $C_p$  reveal that while for high and low  $C_p$  groups the daily variation has individual characteristics, the intermediate  $C_p$  group has no definite type of daily variation associated with it. The study of subgroups of  $C_p$  and I confirm the above conclusions.

e. The amplitudes  $r_1$  and  $r_2$  of diurnal and the semidiurnal component respectively are significant in all cases.  $r_1/r_2$  ratio is important for differentiation of type only for high  $c_p$  days where  $r_1/r_2 \le 1$  is related to  $I^+$  and  $r_1/r_2 \gg 1$  is related to  $I^-$ .

- f. High  $\sum^2$  days represent days of high anisotropy. It days are associated with negative deviations at 6 hour in the mean daily variation curve and I days with a negative deviation at 22 hour. This suggests that both increases as well as decreases of cosmic ray intensity can occur. g. From a comparison of the diurnal and the semidiurnal component of the daily variation of neutrons  $N_k$  and mesons  $M_k$ , it is observed that the diurnal and the semidiurnal components at Kodaikanal are equally important.
- 3. The method of superposed epochs involving what is known as Chree analysis has been used to study the nature of sequential changes of a parameter on days preceding, during and following the occurrence of any particular event. High  $C_p$ , low  $C_p$ , high I and low I events are selected t/ $\phi$  study the sequential changes. The study of I and  $C_p$  for high anisotropy days as indicated by high  $\sum^2$  is also carried out. The following are the main results of the study.
  - a. The decreases of intensity are, in general, connected, to a large extent, with the effect of the beam not explicitly connected with the electric field. However, if the electric effect is also present, high anisotropy occurs.
  - b. An examination of the time series of  $C_p$  for high I and low I indicates the presence of a component of the frozen magnetic field in the direction of the solar dipole field. However, during the period of observation under study the solar dipole field and the sunspot field were in the same direction. Therefore the results obtained cannot be

considered either to confirm or refute the hypothesis of the beams carrying the sunspot fields with them.

- c. It is also observed that the decreases in cosmic ray intensity have a larger amplitude than increases. This is explained by the fact that while decrease can be due to both the electric and magnetic effects of the beam, the increases can be only due to the electric effect.
- 4. With the guidance of Professor H.V.Neher the author has developed an all metal copper system for the purification to a high degree of  $BF_3$  and  $B(CH_3)_3$  gases. The method of purification of  $BF_3$  and the method of preparation of  $BF_3$  counters are described in the second chapter.
- The present dissertation describes in Chapter III the apparatus and in Chapter IV the method of analysis of the investigation made with local neutrons. The principle conclusions of the investigations are discussed in Chapter VI. The author has included at the end of his thesis a list of references to original papers.

Satya Prakash

### Acknowledgments

The work described in this thesis was carried out by the author under the guidence of Professor Vikram A.Sarabhai. I record my deep sense of gratitude to him for the encouragement I received from him throughout the work and the valuable help he gave in interpreting the results. The work on metal system for the preparation of Boron trifluoride counters was carried out with the kind and helpful guidence of Professor H.V.Neher of the California Thatitute of Technology. I have great pleasure in acknowledging my indebtedness for his keen interest. Professor H.A.Ramanathan was evincing a keen interest in the progress of the work and was also guiding me in the project. I wish to record my thanks for his guidance.

I am very thankful to Messrs T.S.G.Sastry and B.A.Holla who were maintaining the neutron pile during the period of study at Kodaikanal. Dr. U.D.Desai and Dr.R.P.Kane were helpful with their discussions from time to time and I am very thankful to them. Messrs R.Sethuraman, S.R.Thakore, H.L.Razdan and Smt. Sushila Rani were helping me in various ways during the progress of the work and the preparation of the thesis. Thanks are also due to the other colleagues of the laboratory for their help and cooperation on all occasions.

To the members of the workshop and the counter section staff I express my thanks for their ready help whenever required.

The project was receiving financial support from the Department of Atomic Energy, Government of India and the author is extremely grateful to them.

# CONTENTS

CHAPTER	I. INTRODUCTION		1	
1.1	Advantages of studying the variation of the cosmic ray nucleonic component	• • •	2	
1.2	The nucleonic component of cosmic rays in the atmosphere	.● @ ● .	3	
1.3	The relationship between the primary cosmic radiation and the observed component of cosmic rays	•••	6	
1.4	The geomagnetic effect in cosmic rays	• • •	8	
1.5	Correction for meteorological elements	• • •	11	
	1.5-1 Correction for pressure changes		12	
	1.5-2 The temperature correction		13	
1.6	Long term changes of intensity of primary radiation	<b>*</b> * 0	18	
	1.6-1 Changes of total intensity with solar cycle	-J) * • •	18	
æ	1.6-2 Long term changes of the knee of the latitude curve	• • •	19	
1.7	27-day variation of cosmic ray intensity	<b>7</b> ••.	21	
1.8	Solar flare type variation of cosmic ray intensity	• • •	25	
1.9	The solar anisotropy of primary radiation	الوز ي رهه.	29	
1.10	Long term changes of solar anisotropy		35	
1.11	Day-to-day changes of the daily variation and their connection with geomagnetic activity			
1.12	The interpretation of the solar anisotropy of cosmic radiation	• • •	. 42	
1.13	Statement of the problem	* * *	50	

CHAPTER	II.	PREPARATION AND PROPERTIES	OF
		BORON TRIFLUORIDE COUNTERS	

2.1	Neutro	n detectors	0 <b>0 0</b>	1
2.2	Theory	of proportional counters	***	Z
	2.2-1	Gas multiplication of proportional counter	• • •	2
	2.2-2	Sensitivity of a counter	8 <b>0</b> 0	$\epsilon$
a	2.2-3	Survival probability of electrons inside a boron trifluoride counter		7
	2.2-4	Life of proportional counters	• •	9
	2.2-5	Determination of the counter background	0 0 0	10
	2.2-6	Significance of counting rate curves	• • •	11
2.3	Theory triflu	of the purification of boron pride	* * *	15
2.4	Prepara count <sub>e</sub> s	ation and filling procedure of	• • •	19
	2.4-1	Assembling counters	• • •	19
4 · •	2.4-2	Description of the system for purification of boron trifluoride	• •	22
	2.4-3	Description of the process of purification	• • •	25
2.5	Results	B. Carrier Marie Control of the Cont	<b>6</b> 5 4	29
CHAPTER '	III.	EXPERIMENTAL ARRANGEMENT FOR THE MEASUREMENT OF LOCAL NEUTRONS	i je pr	
3.1	Introdu	action		1
3.2	The pil	ie		2
	3.2-1	The neutron monitor pile	• • •	2
n .	3.2-2	Characteristics of the pile		7
		nics units	* * *	7
	3.3-1	H.V. Supply		පි

	3.3-2 Low voltage supply	• • •	10
	3.3-3 Pulse amplifier	· ·• • •	11
	3.3-4 Pulse height discriminator	· • • • •	15
	3.3-5 Scalar and recording unit	• • •	17
CHAPTER	IV. METHODS OF ANALYSIS		
4.1	Tabulation and testing of the primary data	• • •	1
4.2	Correction due to barometric pressure	***	بر بر بر
4.3	Determination of the solar daily variation	•••	1
4.4	Fourier analysis	9 4 6	4
4.5	Correlation analysis	Q 8 Q	6
4.6	Chree analysis.	• • •	6
4.7	The variance test		7
4.8	Punch card system of tabulation of data		ç
CHAPTER	V. EXPERIMENTAL RESULTS		
5.1	Introduction	<b>6 6</b> 8 1	1
5.2	Day-to-day changes of intensity		î
5.3	The variance $\sum_{i=1}^{\infty} 2^{i}$ of twelve successive deviations	· ;	Į,
5.4	Individual bihourly deviations and their association with the daily mean intensity		6
5.5	Relationship between the deviations of daily mean intensity and the daily variation	par parti	12
5.6	Relationship of daily variation with Cp	<b>A 4 3</b>	14
	Study of histogram of days for $\emptyset_1$ & $\emptyset_2$		17
5.8	The relationship of the daily variation with I and $C_{\mathbf{p}}$	<b></b>	19
5.9	Relationship of the daily variation with variance of bihourly deviations	• • •	23

5.10	Study of sequential changes of I related to occurrence of days of high and low Cp	# 15 ± 16 ± 16 ± 16 ± 16 ± 16 ± 16 ± 16 ±	28
5.11	Study of sequential changes of Cp related to occurrence of days of high and low I	• • •	28
CHAPTER	VI. DISCUSSION		
6.1	The form of the daily variation of cosmic ray intensity at an equitorial mountain station during 1956-1957	<b>9.6</b> 6	1
6 <b>.2</b>	The electromagnetic state of the interplanegary space.	<b>9 4 6</b>	10
6.3	Conclusion	Ø 6 9	20
REFERENC	DES	<b>9</b> 4 <b>9</b>	į